

Modelling dust in KROME

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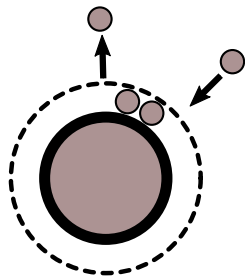
USM/LMU, Munich

November 2018

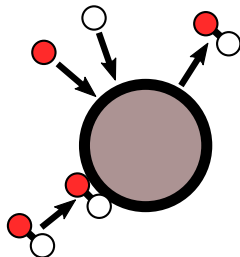


- 1 Recap of dust processes
- 2 Understand dust binning
- 3 how to model dust in KROME
- 4 H_2 formation on dust
- 5 Dust temperature and radiation balance

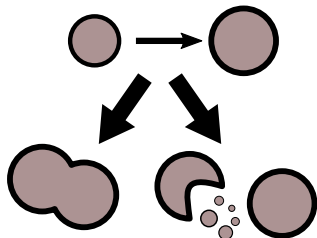
KROME Bootcamp 2018 - Dust process recap



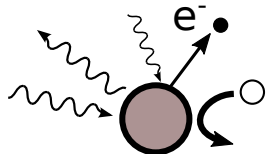
GROWTH
EVAPORATION



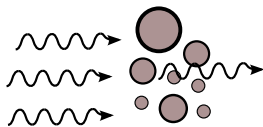
CATALYSIS
FREEZE-OUT



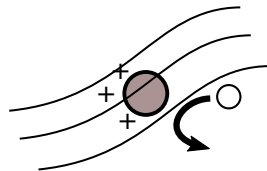
COAGULATION
FRAGMENTATION



COOLING
P.H.E HEATING



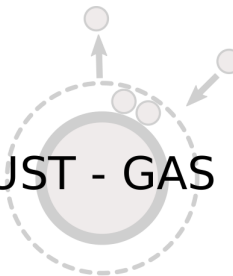
OPACITY



NON-IDEAL MHD

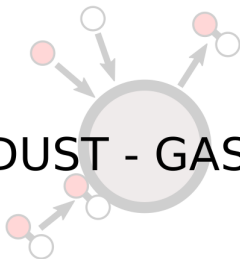
KROME Bootcamp 2018 - Dust process recap

DUST - GAS



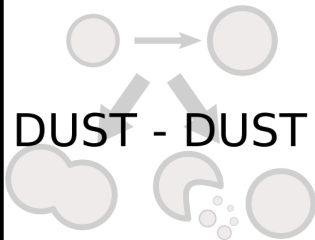
GROWTH
EVAPORATION

DUST - GAS



CATALYSIS
FREEZE-OUT

DUST - DUST



COAGULATION
FRAGMENTATION

DUST - GAS
RADIATION



COOLING
P.E HEATING

DUST
RADIATION

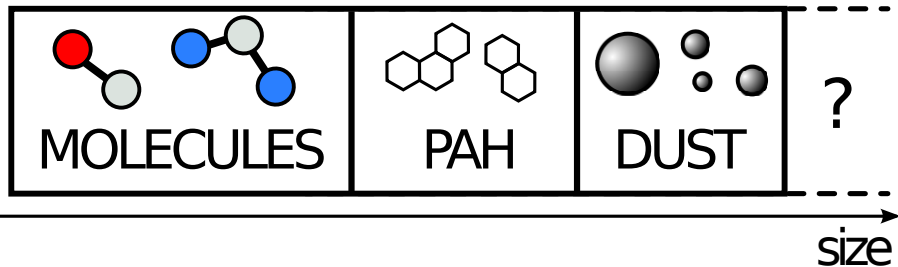


OPACITY

DUST - GAS
B FIELDS



NON-IDEAL MHD



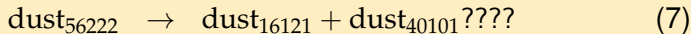
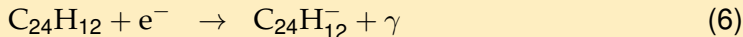
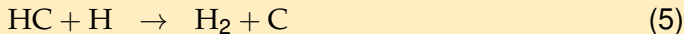
$$\rho_{gr} = 1 \text{ g/cm}^3 \quad (1)$$

$$a = 5 \times 10^{-7} \text{ cm} \quad (2)$$

$$M = \frac{4}{3}\pi a^3 \rho_{gr} \simeq 6 \times 10^{-19} \text{ g} \quad (3)$$

$$N = \frac{M}{m_C} \simeq 3 \times 10^4 \text{ atoms} \quad (4)$$

Reactions for grains



Expand chemical network

H, C, O

H₂, C₂, HC, ...

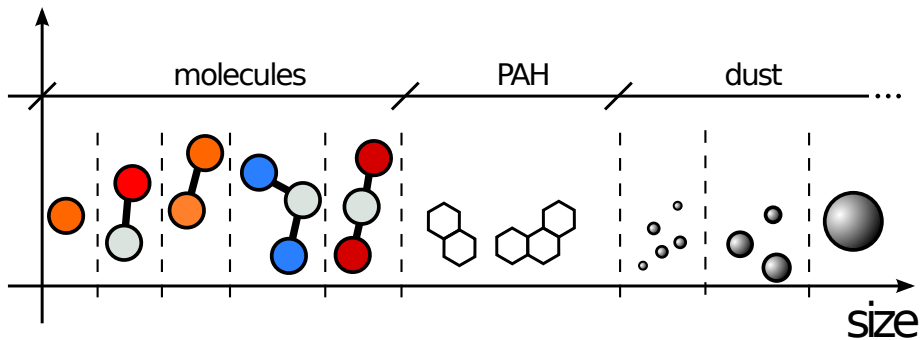
H₂O, CO₂, O₃, ...

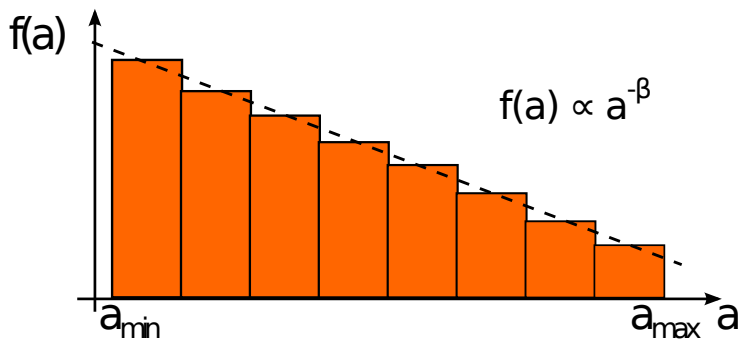
CH₃, H₂O₂, H₂C₂, ...

HC₄, CH₄, H₂CO₂, C₅, C₂H₃, ...

...

Yet another binning...





$$\rho_d = \rho_g D = C \frac{4\pi}{3} \rho_0 \int_{a_{\min}}^{a_{\max}} f(a) a^3 da$$

KROME Bootcamp 2018 - Array size



```
./krome mynetwork.ntw -dust=5,C,Si  
krome_ndust = 5 × 2 = 10  
krome_ndustTypes = 2
```

Initialize the dust

```
call krome_init_dust_distribution(n(:), d2g, &  
                                a_low, a_up, beta)  
call krome_init_dust_distribution(n(:), d2g)
```

Useful dust functions

```
xdust(:) = krome_get_dust_distribution()  
adust(:) = krome_get_dust_size()  
call krome_set_Tdust_array(Tdust(:))  
call krome_get_averaged_Tdust()
```

example

```
./krome -test=dust
```

$$\dot{n}_i = \text{FORMATION} - \text{DESTRUCTION} \pm \text{DUST}_{\text{H} \rightarrow \text{H}_2} [\times N_{\text{gas}}]$$

$$\dot{T} = \text{HEAT} - \text{COOL} \pm \text{DUST} \quad [\times 1]$$

$$\dot{n}_i = \text{COAGULATION} - \text{SPUTTERING} \quad [\times N_{\text{dust}} \times N_{\text{type}}]$$

$$\dot{a}_i = \text{GROWTH} - \text{EVAPORATION} \quad [\times N_{\text{dust}} \times N_{\text{type}}]$$

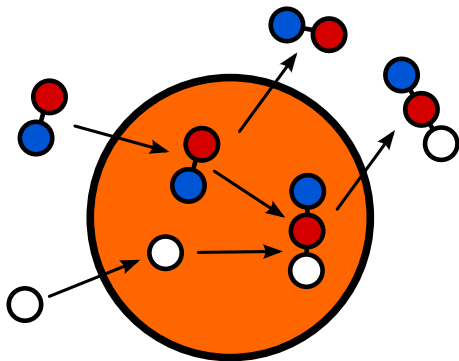
$$\dot{T}_{d,i} = \text{GAS} + \text{RADIATION} \quad [\times N_{\text{dust}} \times N_{\text{type}}]$$

solve a system of stiff ordinary differential equations

H₂ on surface (Cazaux+2009)

$$\dot{n}_{\text{H}_2} = \frac{\pi v_g n_{\text{H}}}{2} \sum_{i \in \text{bins}} \epsilon_i n_{d,i} a_i^2 S_{\text{H}_2}(T_g, T_d)$$

```
./krome -n mynetwork.ntw -dust=10,C -dustOptions=H2
```



Gas-grain chemistry

- Adsorption (gas \rightarrow dust)
- Desorption (dust \rightarrow gas)
- Surface chemistry (dust \rightarrow dust)

Adsorption (Hocuk+2015 and many others)

$X \rightarrow X_{dust}$

$$k_a = \pi a^2 v_g n_d S \quad (8)$$

$$S = \left(1 + 0.4\sqrt{T_g + T_d} + 0.2T_g + 0.08T_g^2\right)^{-1} \quad (9)$$

Desorption (Hocuk+2015 and many others)

$X_{dust} \rightarrow X$

$$k_e = \nu_0 \left[F_{bare} \exp\left(-\frac{E_{bare,i}}{T_d}\right) + F_{ice} \exp\left(-\frac{E_{ice,i}}{T_d}\right) \right] \quad (10)$$

(as it is: likely to be numerically unstable)

Adsorption

$X \rightarrow X_{dust}$

$$k_a = \frac{\pi v_g S \int a^2 \varphi(a) da}{4/3 \pi \rho_0 \int a^3 \varphi(a) da}$$

1, CO, CO_dust, krate_stickSi(n, idx_CO, Tdust)

Desorption

$X_{dust} \rightarrow X$

$$k_e = \nu_0 \exp\left(-\frac{E_i}{T_d}\right)$$

2, CO_dust, CO, krate_evaporation(n, idx_CO, Tdust)

Non-thermal evaporation (Hollenback+2009 and many others)

$X_{dust} + \gamma/CR \rightarrow X$

$$k_{nt,CR} = f_{70} \nu_0 \exp(-E_b/k_B 70 \text{ K})$$

$$f_{70} = 3 \times 10^{-16} \text{ s}^{-1} \frac{\zeta}{\zeta_0}$$

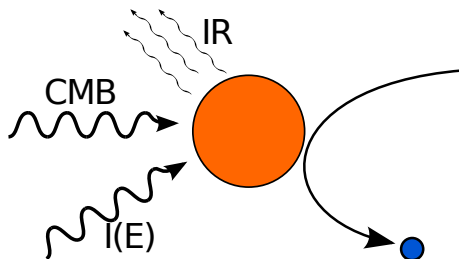
```
3, CO_dust, CO, krate_nonthermal_evaporation  
(idx_CO, Gnot, Av, crflux, yield)
```


Surface chemistry (Hocuk+2015 and many others)



$$k_{ij} = P_{ij} \nu_0 \left[\exp\left(-\delta \frac{E_i}{T_d}\right) + \exp\left(-\delta \frac{E_j}{T_d}\right) \right]$$

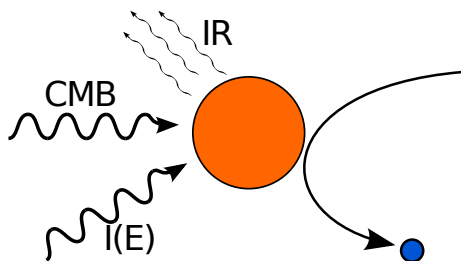
$$P_{ij} = \exp\left[-\frac{2b}{\hbar} \sqrt{2m_{ij} k_B E_a}\right]$$



Kirchoff's law

$$\Gamma_{em} = \Gamma_{abs} \quad (11)$$

$$\int_E Q_{abs} B_{T_d} = \sum_i \int_E Q_{abs} J_i \quad (12)$$



Gas coupling

$$\Gamma_{em} = \Gamma_{abs} + \Lambda_{dust} \quad (13)$$

$$\Lambda_{dust} = 2\pi a^2 n_g n_d v_g k_b (T_g - T_d) \quad \text{erg cm}^{-3} \text{ s}^{-1} \quad (14)$$

Bisection

```
./krome -n mynetwork.ntw -dust=10,C -dustOptions=T  
-cooling=DUST  
solve  $\Gamma_{em}(T_d) = \Gamma_{abs} + \Gamma_{CMB} + \Lambda(T_d, T_g)$  for  $T_d$ 
```

Actual ODE

```
./krome -n mynetwork.ntw -dust=10,C -dustOptions=dT  
-cooling=DUST
```

$$\frac{dT_{d,i}}{dt} = A \frac{dT_g}{dt} \left[A + \int_0^\infty \frac{dB(E, T_{d,i})}{dT_{d,i}} Q_i(E) dE \right]^{-1}$$

$$A = 2 f n_g v_g k_B$$

$$T_{d,i}(t_0 + \Delta t) = T_{d,i}(t_0) + dT_{d,i}/dt$$

(Grassi+2017)

Collapsing cloud - fully consistent

$$[\Gamma_{\text{em}}(T_{d,i}) - \Gamma_{\text{abs}}] \beta_{\text{esc}}(T_{d,i}) = \Lambda(T_{d,i}, T_g) \quad (15)$$

$$\beta_{\text{esc},i} = \tau_{d,i}^{-2} \quad (16)$$

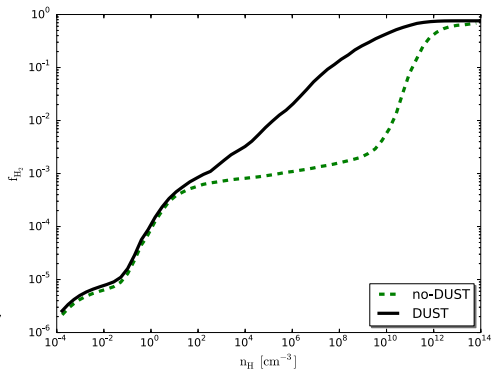
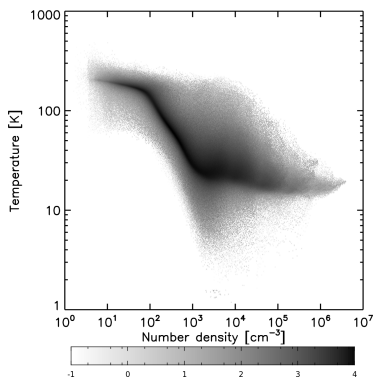
$$\tau_{d,i} = \lambda_{J\pi} \sum_{j \in \text{bins}} n_{d,j} a_j^2 \frac{\int_0^\infty Q(E, a_j) B(E, T_{d,j}) dE}{\int_0^\infty B(E, T_{d,j}) dE} \quad (17)$$

Parameters and table

- dust to gas ratio (metallicity)
- dust distribution (exponent, size range)

$$\begin{aligned} \Lambda_{\text{collapse}} &= f(n_g, T_g) \\ \Lambda_{\text{cloud}} &= f(n_g, T_g, A_V) \end{aligned}$$

KROME Bootcamp 2018 - Dust tables go 3D



What KROME does

- Fully-consistent temperature and cooling/heating (rad-dust-gas)
- Fully-consistent growth and evaporation (dust-gas)
- Fully-consistent H_2 formation (dust-gas)
- Optimized table cooling

What KROME “does”

- Reaction-based surface chemistry (gas-dust)

Whislist

- Fragmentation (dust-dust)
- Coagulation (dust-dust)
- ...

All models are wrong, some are useful

(George E. P. Box)